



Atoms

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NEET: 1-2 Qs (~4-8 marks)

Boards: 4-6 marks



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How we found the nucleus (gold foil), how Bohr quantised the atom to explain hydrogen's line spectrum, and how de Broglie's standing waves justify it all.

1. Rutherford Model & Alpha Scattering

Gold-foil experiment reveals a tiny, massive, positive nucleus in a mostly empty atom.

$$N(\theta) \propto 1/\sin^4(\theta/2)$$

scattering count vs angle

$$r_0 = (1/4\pi\epsilon_0)(2Ze^2/K)$$

distance of closest approach (head-on)

$$b = (Ze^2/4\pi\epsilon_0 K) \cot(\theta/2)$$

impact parameter b vs angle θ

MUST KNOW Most alphas pass straight through (atom is empty); ~1 in 8000 rebound $>90^\circ$ (nucleus is tiny and dense). Nucleus $\sim 10^{-15}$ m, atom $\sim 10^{-10}$ m.

MUST KNOW $r_0 \propto 1/K$ and $\propto 1/m$ — faster (higher-K) alphas approach closer; r_0 is only an UPPER estimate of nuclear size.

- Small $b \rightarrow$ large angle; head-on ($b = 0$) $\rightarrow \theta = 180^\circ$. Drawback: orbiting electron should radiate, spiral in $\rightarrow$ atom unstable + continuous spectrum.

2. Bohr's Postulates

Quantum rules patched onto Rutherford's atom: stable orbits, quantised L , and a frequency condition for radiation.

$$mvr = nh/2\pi \text{ (Bohr quantization)}$$

L quantised, $n = 1, 2, 3, \dots$

$$E = E_2 - E_1 = h\nu$$

frequency / transition condition

MUST KNOW Electrons revolve in stationary orbits WITHOUT radiating, even though they accelerate (centripetal) — this defies classical EM, it is postulated.

MUST KNOW Jump to lower orbit $\rightarrow$ photon EMITTED; jump to higher orbit $\rightarrow$ photon ABSORBED. No radiation while resting in an orbit.

- Works only for hydrogen and hydrogen-like one-electron ions (He^+ , Li^{2+}); fails for multi-electron atoms (even neutral He).

3. Radius, Velocity & Energy of Orbits

Coulomb force = centripetal force, plus the L rule, give clean formulas for each orbit of a hydrogen-like atom.

$$r_n = 0.53 n^2/Z \text{ \AA}$$

$r \propto n^2/Z$; $0.53 \text{ \AA} = \text{Bohr radius}$

$$v_n = 2.18 \times 10^6 Z/n \text{ m/s}$$

$v \propto Z/n$; ground state $\approx c/137$

$$E_n = -13.6 Z^2/n^2 \text{ eV}$$

$E \propto Z^2/n^2$; negative = bound

$$K = -E, U = 2E = -2K$$

virial relations

MUST KNOW Z is in the DENOMINATOR for radius but the NUMERATOR for speed (Z) and energy (Z^2). Don't misplace it.

MUST KNOW H ground-state energy = $-13.6 \text{ eV} \rightarrow$ ionisation energy = 13.6 eV ; for ion, ionisation energy = $13.6 Z^2 \text{ eV}$.

- Ground state $n = 1$, first excited $n = 2$, second excited $n = 3$ — count excited states carefully.

4. Hydrogen Spectrum & Spectral Series

Electron jumps to a lower level n_f emit photons whose wavelength obeys the Rydberg formula; series are named by landing level.

$$1/\lambda = R(1/n_1^2 - 1/n_2^2)$$

Rydberg, $R = 1.097 \times 10^7 \text{ m}^{-1}$, $n_1 < n_2$

$$\text{Number of lines} = n(n-1)/2$$

cascade from level n

MUST KNOW Lyman ($n_f = 1$) UV; Balmer ($n_f = 2$) VISIBLE (only series fully visible); Paschen/Brackett/Pfund ($n_f = 3/4/5$) infrared.

MUST KNOW First line of a series = smallest jump = LONGEST wavelength = least energy; series limit ($n_i = \infty$) = SHORTEST wavelength.

- Lyman series limit = ionisation from ground state (13.6 eV). Keep Z^2 factor for hydrogen-like ions.

5. Limitations & de Broglie Justification

Bohr's quantisation looked arbitrary until de Broglie's standing electron waves explained it.

$$\lambda = h/mv$$

de Broglie wavelength

$$2\pi r = n\lambda \Rightarrow mvr = nh/2\pi$$

standing wave $\rightarrow$ Bohr's L rule

MUST KNOW A stable orbit holds a WHOLE number of electron wavelengths (standing wave); the n th orbit contains exactly n wavelengths.

- de Broglie justifies ONLY the angular-momentum quantisation — not the nucleus, photon emission, or scattering law.
- Bohr's limits: works only for one-electron systems; can't explain fine structure, line intensities, or Zeeman splitting.

Hydrogen Spectral Series

Series	n_f (final)	Transitions	Region
Lyman	1	$n \geq 2 \rightarrow 1$	Ultraviolet
Balmer	2	$n \geq 3 \rightarrow 2$	Visible
Paschen	3	$n \geq 4 \rightarrow 3$	Infrared
Brackett	4	$n \geq 5 \rightarrow 4$	Infrared
Pfund	5	$n \geq 6 \rightarrow 5$	Infrared

n-Dependence (Hydrogen-like)

Quantity	Formula	Varies as
Radius	$0.53 \, n^2/Z \, \text{\AA}$	n^2/Z
Speed	$2.18 \times 10^6 \, Z/n \, \text{m/s}$	Z/n
Energy	$-13.6 \, Z^2/n^2 \, \text{eV}$	Z^2/n^2

Mnemonic - Series order Lyman-Balmer-Paschen-Brackett-Pfund and Lyman=UV, Balmer=Visible, rest=IR

Lyman Born As Baby Pfund — UltraViolet, VISible, then InfraRed

Common traps

- Scattering law is $1/\sin^4(\theta/2)$. A bare $\sin^4(\theta/2)$ option is a deliberate trap.
- $L = nh/2\pi$, NOT nh . Forgetting the 2π is the classic blunder.
- An electron in a stationary orbit IS accelerating (centripetal); Bohr postulates it does not radiate anyway — 'not accelerating' is false.
- Longest wavelength means LEAST energy (smallest jump), not most energetic. Don't swap.
- r_0 is only an upper estimate of nuclear size and only for a head-on hit — not the actual nuclear radius.

Most asked in NEET »

Closest approach $r_0 \propto 1/K, 1/m$

Impact parameter $b = 0 \rightarrow \theta = 180^\circ$

Energy $E_n = -13.6/n^2 \, \text{eV}$

Bohr validity (one-electron only)

Balmer = visible / longest-shortest λ

de Broglie: n wavelengths in n th orbit

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